

**Einstein equations, conformal structure,
and the asymptotic behaviour of space-time**

Cargèse, July 29 – August 10, 2002

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Null cone structure:
main tool for exploration of gravitational fields in the large.

Differential topology and differential geometry

null cone structure \leftrightarrow *causal structure*

causal sets, boundaries of causal sets, horizons, ... ,

Einstein's field equations $\tilde{R}_{\mu\nu} = \lambda \tilde{g}_{\mu\nu} + \kappa (T_{\mu\nu} - \frac{1}{2} T \tilde{g}_{\mu\nu})$

null cone structure \leftrightarrow *characteristics*

positivity conditions, Raychaudhuri equation ... : 'singularities'
formal expansions, analytical control, numerical calculations, ...

Differential geometry and field equations

null cone structure \leftrightarrow *conformal structure*

substructure which is invariant under conformal rescalings

$$\tilde{g} \rightarrow g = \Omega^2 \tilde{g}, \quad \Omega > 0.$$

Ω small: large 'physical' space-time regions are small w.r.t. g -relations.

If $\Omega \rightarrow 0$ at 'infinity' the g -relations may remain finite on a space-time of infinite 'physical' extent

Einstein's equations are not conformally covariant ($n = \dim M$)

$$R_{\nu\rho}[g] = \tilde{R}_{\nu\rho}[\tilde{g}] - \frac{n-2}{\Omega} \nabla_\nu \nabla_\rho \Omega - g_{\nu\rho} \left(\frac{1}{\Omega} \nabla_\lambda \nabla^\lambda \Omega - \frac{n-1}{\Omega^2} \nabla_\lambda \Omega \nabla^\lambda \Omega \right)$$

principal part retained, right hand side degenerates as $\Omega \rightarrow 0$.

Plan

Lecture 1 Formal properties of the field equations,
the conformal field equations.

Lecture 2 The Penrose proposal, problems and results.

Lecture 3 Results on the asymptotic conformal structure
of asymptotically flat vacuum solutions with
vanishing cosmological constant.

Formal properties of the field equations,
the conformal field equations.

Conformal geometry:

A conformal rescaling

$$\tilde{g} \rightarrow g = \Omega^2 \tilde{g}$$

implies transitions $\tilde{\nabla} \rightarrow \nabla$ of the Levi-Civita connection and of the Christoffel symbols

$$\tilde{\Gamma}_{\mu}^{\rho}{}_{\nu} \rightarrow \Gamma_{\mu}^{\rho}{}_{\nu} = \tilde{\Gamma}_{\mu}^{\rho}{}_{\nu} + S(\Omega^{-1} d\Omega)_{\mu}^{\rho}{}_{\nu}$$

where we write, for any 1-form f ,

$$S(f)_{\mu}^{\rho}{}_{\nu} \equiv \delta^{\rho}{}_{\mu} f_{\nu} + \delta^{\rho}{}_{\nu} f_{\mu} - g_{\mu\nu} g^{\rho\lambda} f_{\lambda}.$$

Decomposition of curvature tensor ($\dim M = n \geq 4$)

$$R^{\mu}{}_{\nu\lambda\rho} = 2 \{g^{\mu}{}_{[\lambda} L_{\rho]\nu} - g_{\nu[\lambda} L_{\rho]}{}^{\mu}\} + C^{\mu}{}_{\nu\lambda\rho},$$

conformal Weyl tensor $C^{\mu}{}_{\nu\lambda\rho}$, traces $L_{\mu\nu} \equiv \frac{1}{n-2} \left(R_{\mu\nu} - \frac{1}{2(n-1)} R g_{\mu\nu} \right)$.

Bianchi and once contracted Bianchi identity

$$\nabla_{[\delta} C^{\mu}{}_{|\nu|\lambda\rho]} = 2 (g_{\nu[\lambda} \nabla_{\delta} L_{\rho]}{}^{\mu} - g^{\mu}{}_{[\lambda} \nabla_{\delta} L_{\rho]\nu})$$

$$\nabla_{\mu} C^{\mu}{}_{\nu\lambda\rho} = (n-3) (\nabla_{\lambda} L_{\rho\nu} - \nabla_{\rho} L_{\lambda\nu})$$

Conformal invariance

$$C^{\mu}{}_{\nu\lambda\rho}[g] = \tilde{C}^{\mu}{}_{\nu\lambda\rho}[\tilde{g}]$$

Conformal non-covariance

$$L_{\mu\nu}[g] = \tilde{L}_{\mu\nu}[\tilde{g}] - \frac{1}{\Omega} \nabla_{\mu} \nabla_{\nu} \Omega + \frac{1}{2\Omega^2} g_{\mu\nu} \nabla_{\rho} \Omega \nabla^{\rho} \Omega$$

Conformal covariance

$$\begin{aligned} \nabla_{\mu} (\Omega^{3-n} C^{\mu}{}_{\nu\lambda\rho}) &= \Omega^{3-n} \{ \nabla_{\mu} C^{\mu}{}_{\nu\lambda\rho} - (n-3) \Omega^{n-4} \nabla_{\mu} \Omega (\Omega^{3-n} C^{\mu}{}_{\nu\lambda\rho}) \} \\ &= \Omega^{3-n} \tilde{\nabla}_{\mu} \tilde{C}^{\mu}{}_{\nu\lambda\rho} \end{aligned}$$

The metric conformal field equations I.

Assume that \tilde{g} satisfies Einstein's vacuum field equations $\tilde{R}_{\mu\nu} = \lambda \tilde{g}_{\mu\nu}$ and thus

$$\tilde{L}_{\mu\nu} = \frac{\lambda}{2(n-1)} \tilde{g}_{\mu\nu} \quad \text{and} \quad \tilde{\nabla}_{\mu} C^{\mu}{}_{\nu\lambda\rho} = 0.$$

Exploit available conformal covariance to derive a regular system of equations for Ω and the conformal metric $g = \Omega^2 \tilde{g}$. Set

$$d^{\mu}{}_{\nu\lambda\rho} \equiv \Omega^{3-n} C^{\mu}{}_{\nu\lambda\rho}$$

The equations above imply for the metric and the conformal curvature fields

$$R^{\mu}{}_{\nu\lambda\rho} = 2 \{g^{\mu}{}_{[\lambda} L_{\rho]\nu} - g_{\nu[\lambda} L_{\rho]}{}^{\mu}\} + \Omega^{n-3} d^{\mu}{}_{\nu\lambda\rho}$$

$$\nabla_{\lambda} L_{\rho\nu} - \nabla_{\rho} L_{\lambda\nu} = \Omega^{n-4} \nabla_{\mu} \Omega d^{\mu}{}_{\nu\lambda\rho}$$

$$\nabla_{\mu} d^{\mu}{}_{\nu\lambda\rho} = 0 \quad (\text{Bianchi equation})$$

To obtain equations controlling the evolution of the conformal factor use the trace-free part and the trace of the conformal Ricci tensor

$$\nabla_{\mu} \nabla_{\nu} \Omega = -\Omega L_{\mu\nu} + s g_{\mu\nu},$$

$$\frac{\lambda}{n-1} = 2\Omega s - \nabla_{\rho} \Omega \nabla^{\rho} \Omega.$$

with scalar field

$$s \equiv \frac{1}{n} \nabla_{\lambda} \nabla^{\lambda} \Omega + \frac{1}{2n(n-1)} R \Omega$$

The equations above imply the integrability condition

$$\nabla_{\mu} s = -\nabla^{\nu} \Omega L_{\nu\mu}$$

The metric conformal field equations II.

Unknown

$$u = (g, \Omega, s, L_{\mu\nu}, d^{\mu}{}_{\nu\lambda\rho})$$

Metric conformal field equations

$$R^{\mu}{}_{\nu\lambda\rho} = 2 \{g^{\mu}{}_{[\lambda} L_{\rho]\nu} - g_{\nu[\lambda} L_{\rho]}{}^{\mu}\} + \Omega^{n-3} d^{\mu}{}_{\nu\lambda\rho}$$

$$\nabla_{\mu} \nabla_{\nu} \Omega = -\Omega L_{\mu\nu} + s g_{\mu\nu}$$

$$\nabla_{\mu} s = -\nabla^{\nu} \Omega L_{\nu\mu}$$

$$\nabla_{\lambda} L_{\rho\nu} - \nabla_{\rho} L_{\lambda\nu} = \Omega^{n-4} \nabla_{\mu} \Omega d^{\mu}{}_{\nu\lambda\rho}$$

$$\nabla_{\mu} d^{\mu}{}_{\nu\lambda\rho} = 0$$

$$\frac{\lambda}{n-1} = 2\Omega s - \nabla_{\rho} \Omega \nabla^{\rho} \Omega$$

On a connected set the last equation follows if it holds at one point.

Conformal gauge freedom: equations are invariant under

$$g \rightarrow g^* = \theta^{\frac{4}{n-2}} g, \quad \Omega \rightarrow \Omega^* = \theta^{\frac{2}{n-2}} \Omega, \quad d^{\mu}{}_{\nu\lambda\rho} \rightarrow d^{*\mu}{}_{\nu\lambda\rho} = \theta^{-\frac{2}{n-2}} d^{\mu}{}_{\nu\lambda\rho},$$

$$s \rightarrow s^* = \theta^{-\frac{2}{n-2}} s + \frac{2}{n(n-2)} \theta^{-\frac{2(n-1)}{n-2}} (2\Omega \nabla_{\mu} \theta \nabla^{\mu} \theta + n \theta \nabla_{\mu} \theta \nabla^{\mu} \Omega),$$

$$L_{\mu\nu} \rightarrow L_{\mu\nu}^* = L_{\mu\nu}$$

$$-\frac{2}{(n-2)^2} \theta^{-2} ((n-2) \theta \nabla_{\mu} \nabla_{\nu} \theta - n \nabla_{\mu} \theta \nabla_{\nu} \theta + g_{\mu\nu} \nabla_{\rho} \theta \nabla^{\rho} \theta).$$

In particular

$$L_{\mu}{}^{\mu} \rightarrow L_{\mu}^{*\mu} = \theta^{-\frac{n+2}{n-2}} \left(\theta L_{\mu}{}^{\mu} - \frac{2}{n-2} \nabla_{\mu} \nabla^{\mu} \theta \right).$$

Suggests to remove the gauge freedom by prescribing the function $R = 2(n-1) L_{\mu}{}^{\mu}$ as the conformal gauge source function.

Then Ω is determined implicitly by a wave equation.

The Bianchi equation I.

The vacuum Bianchi identity

$$\tilde{\nabla}_\delta C^\mu{}_{\nu\lambda\rho} + \tilde{\nabla}_\rho C^\mu{}_{\nu\delta\lambda} + \tilde{\nabla}_\lambda C^\mu{}_{\nu\rho\delta} = 0$$

implies in $n \geq 4$ dimensions the system of wave equations

$$\tilde{\nabla}_\delta \tilde{\nabla}^\delta C^\mu{}_{\nu\tau\rho} = 2 \{ C^\mu{}_{\pi\delta\rho} C^\pi{}_{\nu\tau}{}^\delta + C^\mu{}_{\pi\tau\delta} C^\pi{}_{\nu\rho}{}^\delta + C^\mu{}_{\nu\pi\delta} C^\pi{}_{\tau\rho}{}^\delta \} + 2\lambda C^\mu{}_{\nu\tau\rho}$$

and is thus always ‘intrinsically hyperbolic’.

$n = 4$: Bianchi identity \Leftrightarrow contracted Bianchi identity

$$\tilde{\epsilon}_\pi{}^{\delta\tau\rho} \tilde{\nabla}_\delta C^\mu{}_{\nu\tau\rho} = 2 \tilde{\nabla}_\delta C^{*\mu}{}_{\nu\pi}{}^\delta = 2 \tilde{\nabla}_\delta {}^* C^\mu{}_{\nu\pi}{}^\delta = -\tilde{\epsilon}^\mu{}_\nu{}^{\tau\rho} \tilde{\nabla}_\delta C^\delta{}_{\pi\tau\rho}$$

thus: Bianchi equation $\nabla_\delta d^\delta{}_{\pi\tau\rho} = 0$ also ‘intrinsically hyperbolic’.

Conformal Weyl tensor: $u = \frac{1}{12} n^2 (n^2 - 1) - \frac{1}{2} n (n + 1)$ independent components,

Contracted Bianchi id.: $e = \frac{1}{3} n (n - 2) (n + 2)$ independent equations,

$$c = \frac{1}{2} n (n - 1) \text{ constraints}$$

$$(\tilde{\nabla}_\delta C^\delta{}_{0\tau\rho} = 0 \text{ in Gauss coordinates})$$

$n \geq 5$: $e - c < u$ evolution equations in the contracted Bianchi identity.

$\nabla_\delta d^\delta{}_{\pi\tau\rho} = 0$ does not provide hyperbolic evolution equations.

\exists regular, hyperbolic conformal evolution equations if $n \geq 5$?

\exists differences in asymptotic behaviour between $n = 4$ and $n \geq 5$?

Assume from now on: $n = 4$

The Bianchi equation II.

There are various ways to exhibit the hyperbolicity of the Bianchi equation.

If the function t defines a space-like foliation, we write

$$g = (N dt)^2 - h_{ij} (S^i dt + dx^i) (S^j dt + dx^j)$$

$$n^\mu = \frac{1}{N} (\delta^\mu_0 - S^\mu), \quad a_\mu = \frac{1}{N} D_\mu N, \quad h^\mu{}_\nu = g^\mu{}_\nu - n^\mu n_\nu, \quad \chi_{\mu\nu} = \frac{1}{2} \mathcal{L}_n h_{\mu\nu}$$

$$C_{\mu\nu\lambda\rho} = 2 \{ l_{\mu[\lambda} E_{\rho]\nu} - l_{\nu[\lambda} E_{\rho]\mu} - n_{[\lambda} B_{\rho]\tau} \epsilon^\tau{}_{\mu\nu} - n_{[\mu} B_{\nu]\tau} \epsilon^\tau{}_{\lambda\rho} \}$$

and denote by D the h Levi-Civita operator on $\{t = \text{const.}\}$,

we obtain from the Bianchi equations the constraints

$$D^i E_{ij} - B_{ik} \chi^i{}_{l} \epsilon^{kl}{}_j = 0$$

$$D^i B_{ij} + E_{ik} \chi^i{}_{l} \epsilon^{kl}{}_j = 0,$$

which are uniquely determined by the hypersurface $\{t = \text{const.}\}$

and propagation equations

$$\begin{aligned} \frac{1}{N} (\partial_t - \mathcal{L}_S) E_{ij} + D_k B_{l(i} \epsilon_j)^{kl} - 3 E^k{}_{(i} \chi_{j)k} + \chi_k{}^k E_{ij} - \epsilon_i{}^{kl} E_{km} \chi_{ln} \epsilon_j{}^{mn} \\ + 2 a_k B_{l(i} \epsilon_j)^{kl} = 0, \end{aligned}$$

$$\begin{aligned} \frac{1}{N} (\partial_t - \mathcal{L}_S) B_{ij} - D_k E_{l(i} \epsilon_j)^{kl} - 3 B^k{}_{(i} \chi_{j)k} + \chi_k{}^k B_{ij} - \epsilon_i{}^{kl} B_{km} \chi_{ln} \epsilon_j{}^{mn} \\ - 2 a_k E_{l(i} \epsilon_j)^{kl} = 0, \end{aligned}$$

which are not unique.

The latter imply the manifestly symmetric hyperbolic system

$$h^{k(i} h^{j)l} \frac{1}{N} (\partial_t - \mathcal{L}_S) E_{ij} + \frac{1}{2} (\epsilon^{ij(k} h^{l)m} + \epsilon^{im(k} h^{l)j}) D_i B_{jm} + \dots = 0,$$

$$h^{k(i} h^{j)l} \frac{1}{N} (\partial_t - \mathcal{L}_S) B_{ij} - \frac{1}{2} (\epsilon^{ij(k} h^{l)m} + \epsilon^{im(k} h^{l)j}) D_i E_{jm} + \dots = 0.$$

The Bianchi equation III.

In the spin frame formalism $d^\mu{}_{\nu\lambda\rho}$ is represented by a symmetric spinor field ϕ_{abcd}

The Bianchi equation reads

$$0 = \Lambda_{abca'} \equiv \epsilon^{ef} \nabla_{a'f} \phi_{abce}$$

with $a, b, c, \dots, a', b', c' \dots = 0, 1$, and $\epsilon^{ab} = \epsilon^{[ab]}$, $\epsilon^{01} = 1$, $\epsilon^{ab} \epsilon_{ac} = \delta^b{}_c$, $\nabla_{aa'}$ the covariant derivative in the direction of $e_{aa'}$, where

$$e_{00'} = 1/\sqrt{2}(e_0 + e_3), \quad e_{11'} = 1/\sqrt{2}(e_0 - e_3),$$

$$e_{01'} = 1/\sqrt{2}(e_1 + i e_2), \quad e_{10'} = 1/\sqrt{2}(e_1 - i e_2)$$

with a frame field e_k satisfying $g(e_i, e_k) = \eta_{ik}$ such that $g(e_{aa'}, e_{bb'}) = \epsilon_{ab} \epsilon_{a'b'}$.

With $\tau^{aa'} = \delta_0{}^a \delta_{0'}{}^{a'} + \delta_1{}^a \delta_{1'}{}^{a'}$, $P = \tau^{aa'} \nabla_{aa'} = \sqrt{2} \nabla_{e_0}$, $\mathcal{D}_{ab} = \tau_{(a}{}^{a'} \nabla_{b)a'} \perp P$:

$$\Lambda_{abca'} = 0 \quad \text{iff} \quad \Lambda_{abcd} \equiv \Lambda_{abca'} \tau_d{}^{a'} = 0 \quad \text{iff}$$

$$0 = \Lambda_{abf}{}^f = \mathcal{D}^{ef} \phi_{abef} \quad (\text{constraints}), \quad 0 = \Lambda_{(abcd)} = -\frac{1}{2} P \phi_{abcd} + \mathcal{D}_{(d}{}^f \phi_{abc)f}$$

with symmetric hyperbolic version (10 real equations)

$$-\binom{4}{a+b+c+d} \Lambda_{(abcd)} = 0.$$

Further symmetric hyperbolic systems:

$$0 = 2a \Lambda_{(0001')},$$

$$0 = (c-d) \Lambda_{(0011')} - 2a \Lambda_{(0000')},$$

$$0 = (c+d) \Lambda_{(0111')} - (c-d) \Lambda_{(0010')},$$

$$0 = 2e \Lambda_{(1111')} - (c+d) \Lambda_{(0110')},$$

$$0 = -2e \Lambda_{(1110')},$$

with $a, c, e > 0$, $-(2e+c) < d < 2a+c$.

The conformal constraints I.

S space-like hypersurface, $\{e_k\}_{k=0,1,2,3}$ frame near S , $g(e_j, e_k) = \eta_{jk}$, $e_0 \perp S$
Write tensors in this frame, assume $a, b, c, \dots = 1, 2, 3$ and sum convention.
Denote by D the Levi-Civita connection of the induced metric h , set

$$h_{ab} = h(e_a, e_b), \quad \epsilon_{abc} = \epsilon_{[abc]}, \quad \text{with } \epsilon_{123} = 1$$

$$\Sigma = e_0(\Omega), \quad \chi_{ab} = g(\nabla_{c_a} e_0, c_b), \quad L_a = L_{a0}, \quad d_{ab} = d_{a0b0}, \quad d_{ab}^* = d_{a0b0}^*$$

such that

$$d_{ab} = d_{(ab)}, \quad d_a^a = 0, \quad d_{ab}^* = d_{(ab)}^*, \quad d_a^{*a} = 0, \quad d_{abcd} = 2\{h_{a[c}d_{d]b} + h_{b[d}d_{c]a}\}$$

Interior equations implied by metric conformal field equations on S :

$${}^3R_{ab} = \chi_c^c \chi_{ab} - \chi_{ca} \chi_b^c + \Omega d_{ab} + L_{ab} + h_{ab} L_c^c$$

$$D_b \chi_{ca} - D_c \chi_{ba} = \Omega d_{ae}^* \epsilon^e{}_{bc} + 2 h_{a[b} L_{c]}$$

$$D_a D_b \Omega = -\Sigma \chi_{ab} - \Omega L_{ab} + s h_{ab}$$

$$D_a \Sigma = \chi_a^c D_c \Omega - \Omega L_a$$

$$D_a s = -D^b \Omega L_{ba} - \Sigma L_a$$

$$D_a L_{bc} - D_b L_{ac} = D^e \Omega d_{ecab} - \Sigma d_{ce}^* \epsilon^e{}_{ab} + 2 \chi_{c[a} L_{b]}$$

$$D_a L_b - D_b L_a = D^f \Omega d_{fe}^* \epsilon^e{}_{ab} + 2 \chi_{[a}^c L_{b]c}$$

$$D^a d_{ab}^* = -\chi^c{}_e d_{cf} \epsilon_b{}^{ef}$$

$$D^a d_{ab} = \chi^c{}_e d_{cf}^* \epsilon_b{}^{ef}$$

$$\lambda = 6 \Omega s - 3 \Sigma^2 - 3 D_a \Omega D^a \Omega$$

Conformal constraints complicated because of integrability conditions and conformal rescaling. The standard vacuum constraints read

$${}^3R = (\chi_c^c)^2 - \chi_{cd} \chi^{cd} + 2 \lambda, \quad D_a \chi_c^a - D_c \chi_a^a = 0.$$

The conformal constraints II.

Standard vacuum constraints

$${}^3R = (\chi_c{}^c)^2 - \chi_{cd}\chi^{cd} + 2\lambda, \quad D_a\chi_c{}^a - D_c\chi_a{}^a = 0.$$

The conformal field equations include integrability conditions, thus even for $\Omega \equiv 1$ we find the more complicated equations

$${}^3R_{ab} = \chi_c{}^c \chi_{ab} - \chi_{ca}\chi_b{}^c + 2/3 \lambda h_{ab} + d_{ab}$$

$$D_b\chi_{ca} - D_c\chi_{ba} = d_{ae}^* \epsilon^e{}_{bc}$$

$$D^a d_{ab}^* = -\chi^c{}_{e} d_{cf} \epsilon_b{}^{ef}$$

$$D^a d_{ab} = \chi^c{}_{e} d_{cf}^* \epsilon_b{}^{ef}$$

Standard strategy to provide solutions:

Prescribe free data: conformal metric \hat{h}_{ab} , ‘trial’ $\hat{\chi}_{ab}$,

Solve linear resp. semi-linear elliptic equations to obtain solution to the standard constraints (Lichnerowicz, York, Choquet-Bruhat, ...)

Use first subsystem above to define d_{ab} and d_{ab}^* .

The other equations will then be satisfied.

Procedure works for conformal constraints similarly.

Delicate: behaviour near $\Omega = 0$.

Different strategy:

Prescribe free data: ‘trial’ \hat{d}_{ab} , \hat{d}_{ab}^*

Read the system as a 3-dimensional Einstein equation with source fields

$$d_{ab} = D_a X_b + D_b X_a - \frac{2}{3} h_{ab} D_c X^c + \hat{d}_{ab}$$

$$d_{ab}^* = D_a X_b^* + D_b X_a^* - \frac{2}{3} h_{ab} D_c X^{*c} + \hat{d}_{ab}^*$$

Solve quasi-linear elliptic system for h_{ab} , χ_{ab} , X_a , X_a^* .

First task: formulate well posed elliptic boundary value problems.

A. Butscher (2002): Stability of asymptotically euclidean vacuum solution $h_{ab} = -\delta_{ab}$, $\chi_{ab} = 0$, $d_{ab} = 0$, $d_{ab}^* = 0$.

Gauge conditions, hyperbolicity, and local evolution.

The conformal field equations pose the same problems and offer the same flexibility as Einstein's equations in their standard form.

∃ various ways to handle conformal, coordinate, and frame gauge freedom:

Implicit wave equations for conformal factor, coordinates, frame field, or geometrical gauge with explicit algebraic conditions on unknown, ...

All techniques developed for Einstein's field equations may be tried:

Writing formally $\kappa T_{\mu\nu} \equiv 2(L_{\mu\nu} - L g_{\mu\nu})$, the metric conformal field equations take the form of

‘Einstein equations with matter fields’

$$\begin{aligned} R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} &= \kappa T_{\mu\nu} \\ \nabla_{\mu} \nabla_{\nu} \Omega &= -\frac{\kappa}{2} (T_{\mu\nu} - \frac{1}{3} T g_{\mu\nu}) \Omega + s g_{\mu\nu} \\ \nabla_{\mu} s &= -\frac{\kappa}{2} (T_{\mu\nu} - \frac{1}{3} T g_{\mu\nu}) \nabla^{\nu} \Omega \\ \kappa (\nabla_{[\lambda} T_{\rho]\nu} - \frac{1}{3} \nabla_{[\lambda} T g_{\rho]\nu}) &= \nabla_{\mu} \Omega d^{\mu}{}_{\nu\lambda\rho} \\ \nabla_{\mu} d^{\mu}{}_{\nu\lambda\rho} &= 0 \end{aligned}$$

(ensure geometrical role of $T_{\mu\nu}$ and $d^{\mu}{}_{\nu\lambda\rho}$)

∃ hyperbolic systems in the tensor, frame, and spin frame formalism.

∃ various ways to formulate well posed local initial value problems for the metric conformal field equations (preserving constraints and gauge conditions)

Specific for the conformal field equations:

Various global and semi-global geometric problems for Einstein's field equations can be/have been recast as well posed

- characteristics initial value problems,
- Cauchy problems,
- initial boundary value problems

for the conformal field equations, involving finite domains where Ω changes sign.

The general conformal field equations I.

Besides conformal rescalings $\tilde{g} \rightarrow g = \Theta^2 \tilde{g}$

consider also Weyl connections $\hat{\nabla} = \nabla_g + S(f)$ with arbitrary 1-form f , torsion free, respect conformal class \mathcal{C}_g : $\hat{\nabla}_\rho g_{\mu\nu} = -2 f_\rho g_{\mu\nu}$.

Representation of 1-form depends on conformal scaling of metric

$$\nabla_g = \nabla_{\tilde{g}} + S(\Theta^{-1} \nabla \Theta), \quad \hat{\nabla} = \nabla_{\tilde{g}} + S(\Theta^{-1} d) \quad \text{with } d = \Theta f + \nabla \Theta.$$

Write fields in frame $\{e_k\}_{k=0,\dots,3}$ with $g(e_i, e_k) = \eta_{ik}$. Set

$$\hat{\nabla}_i e_j \equiv \hat{\nabla}_{e_i} e_j = \hat{\Gamma}_i^k{}_j e_k, \quad d^i{}_{jkl} = \Theta^{-1} C^i{}_{jkl},$$

$$\hat{L}_{jk} = 1/2 \hat{R}_{(jk)} - 1/4 \hat{R}_{[jk]} - 1/12 \hat{R} g_{jk}$$

The vacuum field equation $Ric[\tilde{g}] = \lambda \tilde{g}$ then imply for

$$u = (e^\mu{}_k, \hat{\Gamma}_i^j{}_k, \hat{L}_{jk}, d^i{}_{jkl}),$$

the ‘conformal equations based on Weyl connections’

$$[e_p, e_q] = (\hat{\Gamma}_p^l{}_q - \hat{\Gamma}_q^l{}_p) e_l,$$

$$\begin{aligned} e_p(\hat{\Gamma}_q^i{}_j) - e_q(\hat{\Gamma}_p^i{}_j) - \hat{\Gamma}_k^i{}_j (\hat{\Gamma}_p^k{}_q - \hat{\Gamma}_q^k{}_p) + \hat{\Gamma}_p^i{}_k \hat{\Gamma}_q^k{}_j - \hat{\Gamma}_q^i{}_k \hat{\Gamma}_p^k{}_j \\ = 2 \{g^i{}_{[p} \hat{L}_{q]j} - g^i{}_j \hat{L}_{[pq]} - g_{j[p} \hat{L}_{q]}{}^i\} + \Theta d^i{}_{j pq}, \end{aligned}$$

$$\hat{\nabla}_p \hat{L}_{qj} - \hat{\nabla}_q \hat{L}_{pj} = d_i d^i{}_{j pq},$$

$$\nabla_i d^i{}_{jkl} = 0.$$

In the Bianchi equation the connection $\nabla = \nabla_g$ is used, it holds

$$\hat{\Gamma}_i^j{}_k = \Gamma_i^j{}_k + \delta^j{}_i f_k + \delta^j{}_k f_i - \eta_{ik} \eta^{jl} f_l \quad \text{with } f_i = \frac{1}{n} \hat{\Gamma}_i^k{}_k.$$

New gauge freedom. No equations given for Θ and d_i .

Conformal geodesics.

Conformal geodesic:

curve $x(\tau)$ with 1-form $b(\tau)$ along $x(\tau)$ such that

$$(*_1) \quad (\tilde{\nabla}_{\dot{x}} \dot{x})^\mu + S(b)_\lambda{}^\mu{}_\rho \dot{x}^\lambda \dot{x}^\rho = 0,$$

$$(*_2) \quad (\tilde{\nabla}_{\dot{x}} b)_\nu - \frac{1}{2} b_\mu S(b)_\lambda{}^\mu{}_\nu \dot{x}^\lambda = \tilde{L}_{\lambda\nu} \dot{x}^\lambda,$$

where $\tilde{L}_{\mu\nu} = \frac{1}{2} (\tilde{R}_{\mu\nu} - \frac{1}{6} \tilde{R} \tilde{g}_{\mu\nu})$,

$x(\tau)$ is a conformal invariant: it is still solution after $\tilde{g} \rightarrow \Omega^2 \tilde{g}$, $\tilde{\nabla} \rightarrow \hat{\nabla}$,

Conformal geodesic specified by data $x(\tau_*) = x_*$, $\dot{x}(\tau_*) = \dot{x}_*$, $b(\tau_*) = b_*$.

sign of $\tilde{g}(\dot{x}, \dot{x})$ preserved along $x(\tau)$: $\tilde{\nabla}_{\dot{x}}(\tilde{g}(\dot{x}, \dot{x})) = -2 \langle b, \dot{x} \rangle \tilde{g}(\dot{x}, \dot{x})$,

$x(\tau)$, $b(\tau)$ conf. geod. then $\bar{x}(\bar{\tau}) = x(\tau(\bar{\tau}))$, $\bar{b}(\bar{\tau})$ conf. geod, iff

$$\tau = \tau_* + \frac{\bar{\tau} - \bar{\tau}_*}{e + c(\bar{\tau} - \bar{\tau}_*)}, \quad \bar{b} = b + \frac{1}{\tilde{g}(\dot{x}, \dot{x})} \frac{2c}{1 - c(\bar{\tau} - \tau_*)} \dot{x}_b, \quad e, c, \tau_*, \bar{\tau}_* \in \mathbb{R}, \quad e \neq 0.$$

(*3) $(\tilde{\nabla}_{\dot{x}} e_k)^\mu + S(b)_\lambda{}^\mu{}_\rho \dot{x}^\lambda e^\rho{}_k = 0$ respects $\tilde{g}(e_i, e_k) = \Theta^2 \eta_{ik}$ with

(*4) $\tilde{\nabla}_{\dot{x}} \Theta = \Theta \langle b, \dot{x} \rangle$, $\Theta_* > 0$.

Conformal Gauss gauge (including x^μ , e_k , Θ) based on space-like hypersurface S :

choose coordinates x^α on S , $\alpha = 1, 2, 3$,

prescribe smooth data x_* , \dot{x}_* , b_* , $\Theta_* > 0$, e_{k*} on S with:

$$\dot{x}_* = e_{0*} \perp S, \quad \tilde{g}(e_i, e_k)_* = \Theta_*^2 \eta_{ik},$$

solve (*1), (*2), (*3), (*4) and $\tilde{\nabla}_{\dot{x}} x^\alpha = 0$,

set $x^0 = \tau$, $S = \{\tau = 0\}$.

The slices $\{\tau = \text{const.} > 0\}$ are not necessarily orthogonal to $\dot{x} = e_0$!

The data on S are given ‘by hand’, the rest, in particular the conformal scaling, is determined by the conformal structure.

Conformal geodesics on vacuum solutions.

If Einstein's equations $Ric[\tilde{g}] = \lambda \tilde{g}$ hold and $\tilde{g}(\dot{x}, \dot{x}) > 0$, then

Scaling and parametrization:

Θ and $d_k = \Theta b_\mu e^\mu{}_k$ can be integrated along the conformal geodesics. The explicit expressions for Θ and d_k are

$$\Theta = \Theta_* \left(1 + \tau \langle b_*, \dot{x}_* \rangle + \frac{\tau^2}{2} \left(\Theta_*^{-2} \frac{\lambda}{6} + \frac{1}{2} g^\sharp(b_*, b_*) \right) \right),$$
$$d_0 = \dot{\Theta}, \quad d_a = \langle b_*, \Theta_* e_{a*} \rangle, \quad a = 1, 2, 3,$$

These expressions encode information on the field equations, in particular (on a smooth, nondegenerate congruence):

$$\nabla_k \Theta \nabla^k \Theta = -\frac{1}{3} \lambda \quad \text{where} \quad \Theta = 0.$$

For suitable data the expression for Θ admits 2 zeros.

Point sets:

Reparametrize $x(\tau)$ such that $\bar{x}(\bar{\tau}) = x(\tau(\bar{\tau}))$ satisfies $\tilde{g}(\bar{x}', \bar{x}') = 1$.

Write $b = \hat{b} + \zeta \dot{x}^b$ such that $\langle \hat{b}, \dot{x} \rangle = 0$. It follows:

on vacuum fields the conformal geodesic equations imply the 'vacuum-adapted conformal geodesic equations'

$$\tilde{\nabla}_{\bar{x}'} \bar{x}' = \hat{b}^\sharp, \quad \tilde{\nabla}_{\bar{x}'} \hat{b} = \beta^2 \bar{x}'^b,$$

where $\beta^2 \equiv -\tilde{g}^\sharp(\hat{b}, \hat{b}) = (\delta^{ab} d_a d_b)_* = \text{const. along } \bar{x}(\bar{\tau})$,

Thus:

if \hat{b} vanishes at a point, the curve is a \tilde{g} -geodesic,

the choice of ζ_* determines the parameterization (slicing),

after a reparameterization conformal geodesics satisfy

an equation of third order. Here $\tilde{\nabla}_{\bar{x}'}^2 \bar{x}' = \beta^2 \bar{x}$.

The general conformal field equations II.

We call the ‘conformal equations based on Weyl connections’ in a conformal Gauss gauge the ‘general conformal field equations’.

With the ‘vacuum expressions for Θ and d_k ’ (or with the equations governing Θ and d_k) the equations represent a complete system for

$$u = (e^\mu{}_k, \hat{\Gamma}_i{}^j{}_k, \hat{L}_{jk}, d^i{}_{jkl}).$$

A conformal Gauss gauge implies the explicit conditions

$$\dot{x} = e_0 = \partial_\tau, \quad \hat{\Gamma}_0{}^j{}_k = 0, \quad \hat{L}_{0k} = 0.$$

The general conformal field equations thus imply evolution equations of the form

$$\partial_\tau e^\mu{}_k = -\hat{\Gamma}_k{}^l{}_0 e^\mu{}_l,$$

$$\partial_\tau \hat{\Gamma}_l{}^i{}_j = -\hat{\Gamma}_k{}^i{}_j \hat{\Gamma}_l{}^k{}_0 + g^i{}_0 \hat{L}_{lj} - g_{j0} \hat{L}_l{}^i + g^i{}_j \hat{L}_{l0} + \Theta d^i{}_{j0l},$$

$$\partial_\tau \hat{L}_{kj} + \hat{\Gamma}_k{}^i{}_0 \hat{L}_{ij} = d_i d^i{}_{j0k},$$

$$\nabla_i d^i{}_{jkl} = 0.$$

This system implies symmetric hyperbolic ‘reduced equations’ which preserve the constraints in domains where $\Theta > 0$.

In the conformal Gauss gauge the location of the set $\{\Theta = 0\}$ can be prescribed explicitly in terms of the data.